

NPTEL ONLINE CERTIFICATION COURSES

EARTHQUAKE SEISMOLOGY

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Module 04 : Normal modes of the Earth Lecture 03: Solving numerical problems

CONCEPTS COVERED

➢ Recap

> Solving numerical problems



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Normal modes are "Oscillation of the system" and is sum of the standing waves. Example, whole earth vibration after a large earthquake, free oscillation of sun. For displacement in 3D:

 $\mathbf{u}(r, \theta, \phi) = \sum_{n} \sum_{l} \sum_{m} {}_{n} A_{l}^{m} {}_{n} y_{l}(r) \mathbf{x}_{l}^{m}(\theta, \phi) e^{i_{n} \omega_{l}^{m} t}$ Displacement

n,l, m-radial, angular and azimuthal order

"y_l(r)-scaler radial eigenfunction

 $\mathbf{x}_l^m(heta,\phi)$ - vector surface eigenfunction

 $_{n}A^{m}_{l}$ excitation amplitudes (weights for eigenfunctions) that depend on the seismic source



Two types of Normal modes: Spheroidal modes; analogous to P-SV motion.

Toroidal Modes; analogous to love of SH motion.

Boundary condition on the bounded string (at x=0 and x=L) produces normal modes, and solution is given by

$$Y(x,t) = \sum_{n=0}^{\infty} A_n Y_n(x,\omega_n) \cos(\omega_n t) \qquad Y_n(x,\omega_n) = \sin(\omega_n x/v)$$

* Similarly, fully normalized and orthogonal spherical harmonics for earth is given by $Y_l^m(\theta,\phi) = (-1)^m \left[\left(\frac{2l+1}{4\pi} \right) \frac{(l-m)!}{(l+m)!} \right]^{1/2} P_l^m(\cos\theta) e^{im\phi}$

Where n, I and m are radial, angular and azimuthal order.



Surface eigenfunctions for Toroidal modes are given by a vector spherical harmonics:

$$\mathrm{T}_l^m(r, heta,\,\phi) = \left(0,\, rac{1}{\sin heta} rac{\partial Y_l^m(heta,\,\phi)}{\partial\phi},\, - rac{\partial Y_l^m(heta,\,\phi)}{\partial heta}
ight)$$

The displacements are given by

$$\mathbf{u}^{T}(r,\theta,\phi) = \sum_{n} \sum_{l} \sum_{m=-l}^{l} \underbrace{A_{l}^{m}}_{amplitude} \underbrace{W_{l}(r)}_{radial \ eigenfunction} \underbrace{W_{l}(r)}_{\theta,\phi \ motion} e^{i_{n}\omega_{l}^{m}t}$$



For spheroidal mode, the surface eigenfunctions are given by two other vector spherical harmonics with (r, θ , ϕ) components

$$egin{aligned} \mathbf{R}_l^m &= (Y_l^m, 0, \, 0) \ \mathbf{S}_l^m &= \left(0, \, rac{\partial Y_l^m(heta, \, \phi)}{\partial heta}, \, rac{1}{\sin heta} rac{\partial Y_l^m(heta, \, \phi)}{\partial \phi}
ight) \end{aligned}$$

Each corresponds to a different radial eigenfunction, ${}_{n}U_{l}(r)$ and ${}_{n}V_{l}(r)$, so the displacement for spheroidal modes is

$$\mathbf{u}^{s}(r,\, heta,\,\phi) = \sum_{n} \sum_{l} \sum_{m=-l}^{l} {}_{n}A_{l}^{m} \left[{}_{n}U_{l}(r)R_{l}^{m}(heta,\,\phi) + {}_{n}V_{l}(r)\mathbf{S}_{l}^{m}(heta,\,\phi)
ight] e^{n^{\omega_{l}^{m_{t}}}}$$

The radial eigenfunction $_{n}U_{l}(r)$ corresponds to radial motion and $_{n}V_{l}(r)$ corresponds to horizontal motion.



Problem 1. For m =1, l =2

- a) What is $P_1^m(\cos\theta)$?
- a) What is $Y_2^1(heta, \phi)$?
- a) What is T_2^1 ?

Problem 2. Calculate surface eigenfunctions R_2^2 , S_2^2 , and T_2^2 ?









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