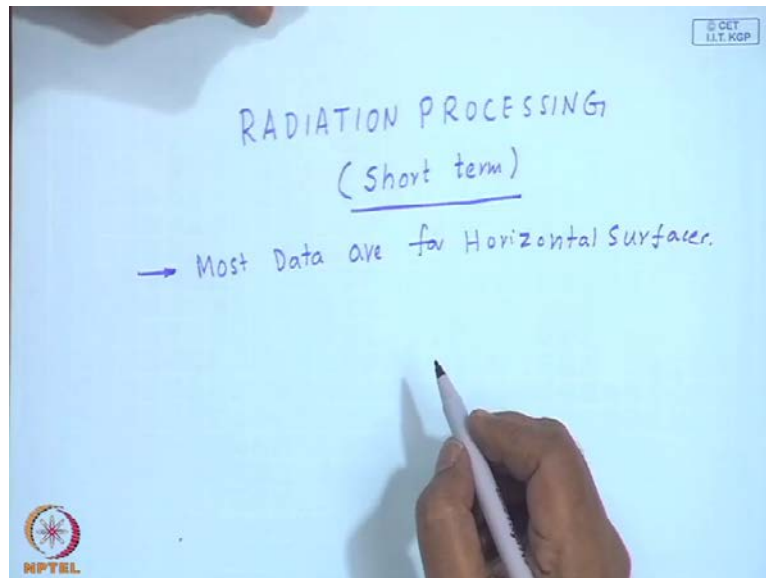


Solar Energy Technology
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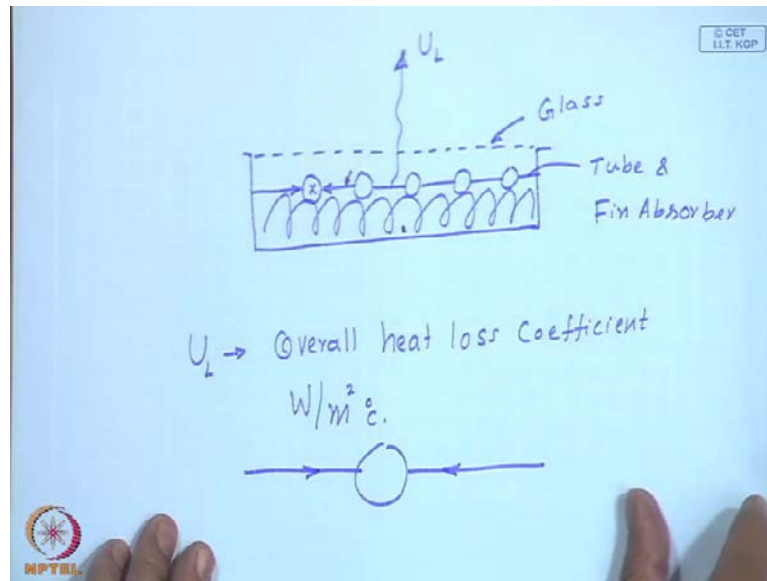
Lecture - 5
Estimation of Solar Radiation or Details

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Now, we learnt about Radiation solar radiation measurement. And some useful correlations, which will help us generating the detail of solar radiation given certain broad averages or long time periods, longer time period. Then, what we shall be discussing here is radiation processing. First I will call it short term, later on we shall go to the long term processing. And we know from the data most data are for horizontal surfaces. Now here, in order that you appreciate why we do this radiation processing, very briefly I shall explain about a simple solar collector. So, that you will understand why we do this particular part.

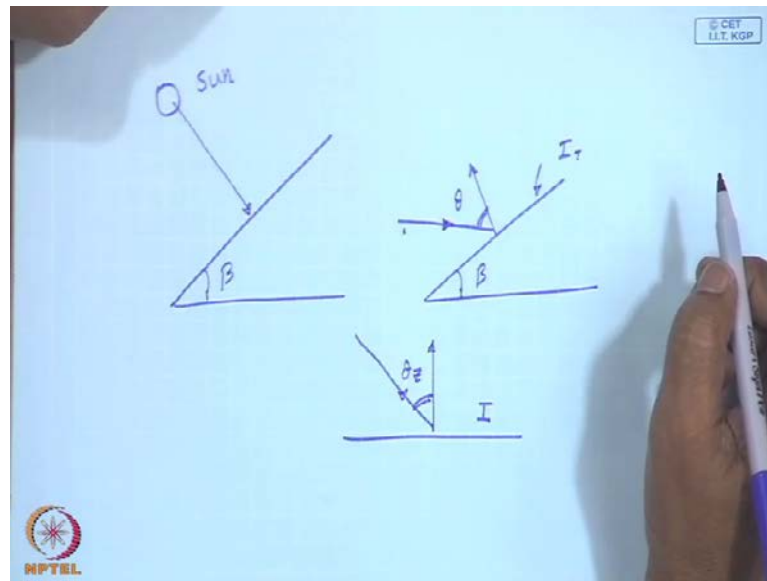
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So, you may have a fin a tube a fin another tube another tube like this we set spacing. This is kept in a box and the bottom of which is insulated. And there is glass cover this is a typical water heater this is called a tube, and fin absorber. So, overall there will be some loss which I will represent by a loss coefficient U_L . So, those are familiar with heat transfer this is called a overall heat loss coefficient. And which will have the units watts per meter square degree centigrade. And the fluid flows through these tubes. The central portion painted black over here. Heat will flow towards the colder fluid side thereby heating the fluid. In other words if, I exaggerate there will be heat flow towards the tube right.

So, this is based upon the fact given per say a black painted surface will absorb solar radiation more than a non-bright non-black painted absorber. And to reduce the losses we have put a insulation at the bottom. And then a glass cover at a top to reduce by radiation, and convection process, still there is a certain loss taking place at the rate of U_L watts per meter square degree centigrade.

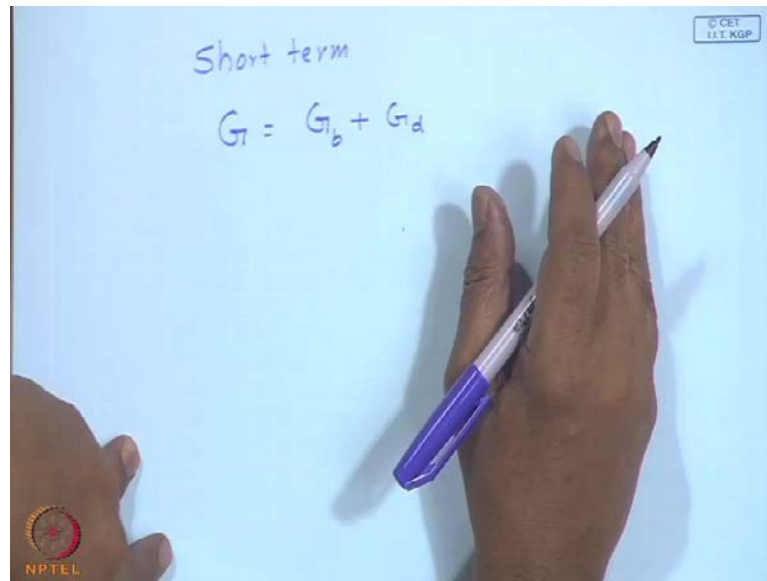
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We, also know that a surface having a particular slope beta will receive maximum radiation if the sun's rays or normal to the surface. If you want to keep it you have to keep on tracking it. But, never the less I may choose some sort of a slope beta with may have may or may not be this is the outer normal to the surface sun's ray may be like this. And the angle of incidence be theta. What I have got is a measurement sun a horizontal surface with the sun's rays making an angle theta z. This is my eye this is my eye on the tilted surface.

So, we want to estimate in general on any surface given the value on the horizontal surface this may be on the hourly time scale. So, our objective will be to find out what is a solar radiation received by a surface of given orientation. Known that there is a solar radiation available measured or otherwise on a horizontal surface.

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So, first since I called it short term I will talk about, in tensed G . which will consist of G_b direct part plus G_d the diffuse part. Now, these are may be divided by 3600 or the solar radiation for the hourly value or let us assume that we know the intensity.

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G_b and G_d are the beam and diffuse components of intensity.
The total radiation tilted surface G_T can be formally expressed as,

$$G_T = G_b R_b + F_s G_d + F_g \rho_g G$$

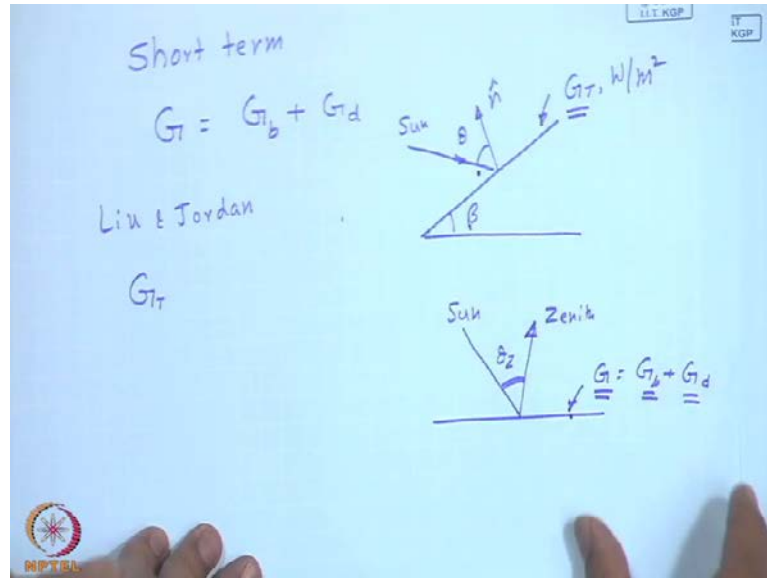
ρ_g is the ground reflectivity (usually 0.2 and 0.7 if the ground is snow covered) and R_b is the instantaneous tilt factor for beam radiation, defined as the the ratio of



First if I know how to do it for a instant then we will be able to find out for this number of developments by Liu and Jordan. Again, they classified that the radiation received by the tilted surface. Now, what I have got is beta outer normal. And the sun's ray making angle theta compare and this is G_t watts per meter square compared to horizontal surface

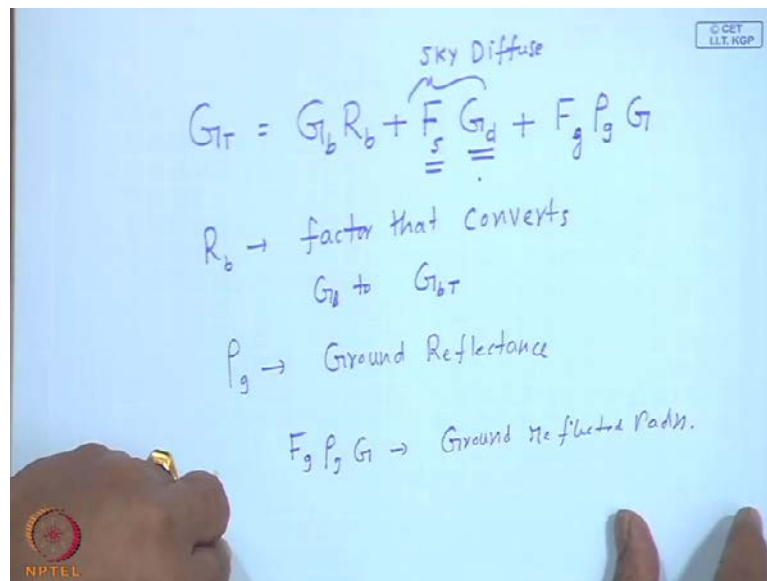
making an angle theta z where, this is G equal to Gb plus Gd o, we have a tilted surface with a slope beta other descriptions right now are unnecessary which we will come to it later whose outer normal is n cap and the sun's rays will make an angle theta in general.

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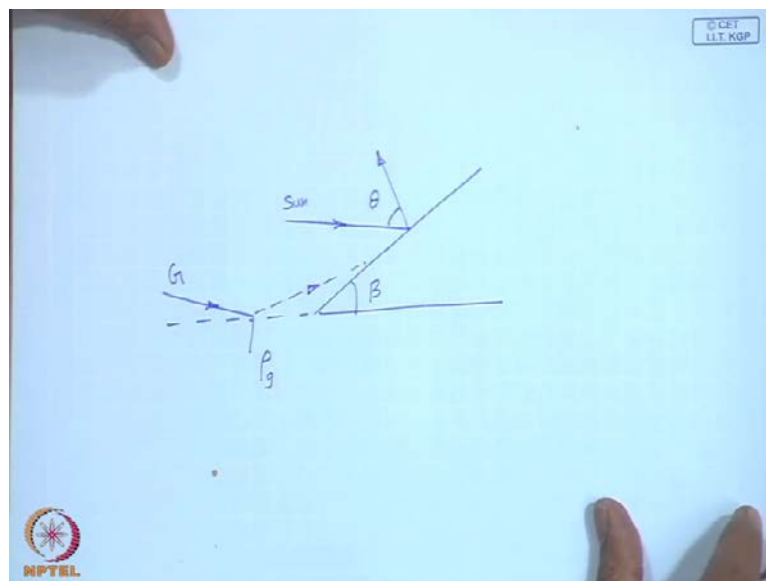
And the radiation falling on that is Gt no matter how to calculate that, what we have is measured information a g intensity on the horizontal plane which is characterized by the zenith, and the sun makes an angle theta z with respect to this we also have the information that G comprises of Gb and Gd.

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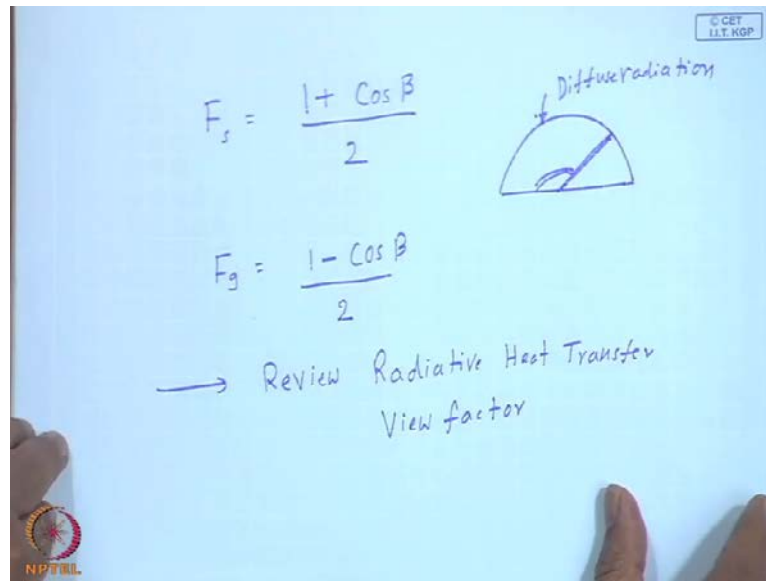
Pupil have expressed this G_t comprises of direct radiation part plus a fraction of the diffuse radiation plus another fraction of the ground effected radiation. So, R_b is a factor that converts G_b to if I may say G_b on the tilted plane, this is purely geometric then G_d is the diffuse part but, a fraction will be received by that which will be called sky diffuse. And ρ_g ground reflectance, so, this F_{gs} , fraction f this should be $F_g \rho_g$ G is the ground reflected radiation.

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Now, to make you appreciate if there is a collector which set and slope again and then of course, I will keep on showing this, where sun's ray with an angle theta is this is sun. And if this is the ground sun's ray will fall on this, this 2 should be parallel. And which will be reflected this is ρ_g the ground reflectance makes G reflect back some of it reflect on to the collector surface, which will give you that $F_g \rho_g G$ into G term.

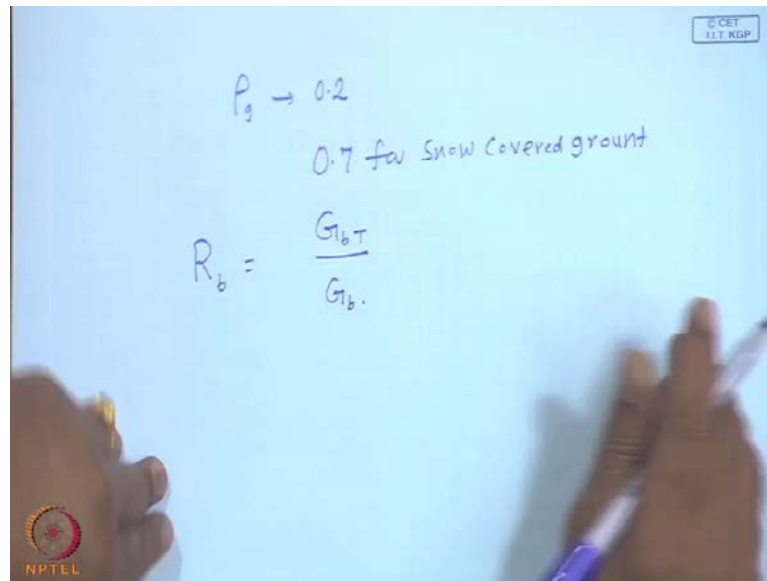
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Similarly, this F_s associated with G_d is given by $1 + \cos \beta$ by 2 where β is a slope. Now, if you imagine a hemisphere it is a collector. All the diffuse radiation over the hemisphere is your G_d . So, this portion will be received by the solar collector. You will have this factor $1 + \cos \beta$ by 2 and if, you make β equal to 0 this $1 + 1$ by 2 by 2 equal to 1 that means if it is a horizontal surface. It will receive in full the diffuse radiation. And again if you have a vertical surface β is equal to 90 degrees making this as one and half. So, a quadrant of that will make it receive only one and half so, sufficient to understand this $1 + \cos \beta$ by 2 comes as a factor associated with the diffuse radiation which basically is a sector proportional assuming the diffused radiation to be isotropic.

Then F_g is again $1 - \cos \beta$ by 2 you should review radiative heat transfer where view factors are given. So, you can see if there are 2 surfaces with a angular orientation the ground effected radiation will be $1 - \cos \beta$ by 2 factor. Again if you put β equal to 0 it will be 0 in other words say horizontal surface will not receive any reflected radiation from ground also be in horizontal.

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beam radiation on the surface under consideration and the beam radiation on a horizontal surface. Thus,

$$R_b = G_{bT}/G_b$$

F_s and F_g are the appropriate factors for sky diffuse and ground reflected components of radiation. Assuming an isotropic distribution of diffuse radiation, the factor F_s can be obtained as,

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And this rho g where is typically point 2 for a ground and it could be as say as point 7 for snow covered. Of course, you may wonder if snow cover is there the solar radiation will be low whether your solar collector should be operating or not. Now this Rb is the instantaneous tilt factor that is what we try to find out.

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$$F_s = (1 + \cos \beta)/2$$

The view factor F_g is the view factor between the ground and the collector

$$F_g = (1 - \cos \beta)/2 \quad (7.5)$$

So, I can write r_b as g_b on the tilted plane by G_b on the horizontal plane and we have already talked about F_s and F_g or the factors associated with the sky diffuse and ground reflected factors. Which are given by $1 + \cos \beta$ by 2 and $1 - \cos \beta$ by 2.

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Expression for R_b

(a) Horizontal Surface

(b) Tilted Surface

Beam radiation on horizontal and tilted surfaces

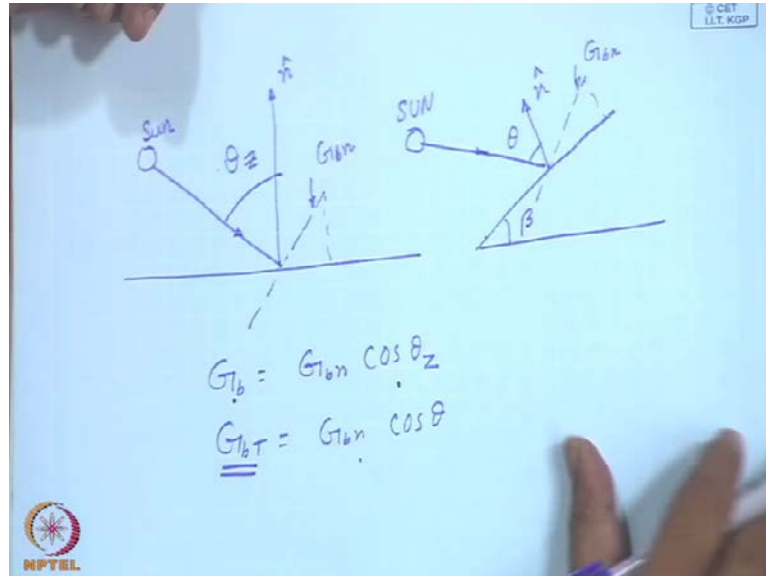
$$G_b = G_{b,n} \cos \theta_z, \quad G_{b,T} = G_{b,n} \cos \theta$$

Thus, R_b can be expressed as,

Now, this figure I should make it little bigger, if you have the horizontal surface. And the outer normal, and this is the sun's ray and it has got the angle θ_z . And simultaneously we have the tilted surface β outer normal. And the sun ray, and this makes an angle θ . So, we can say G_b what I shall do is if I draw a plane

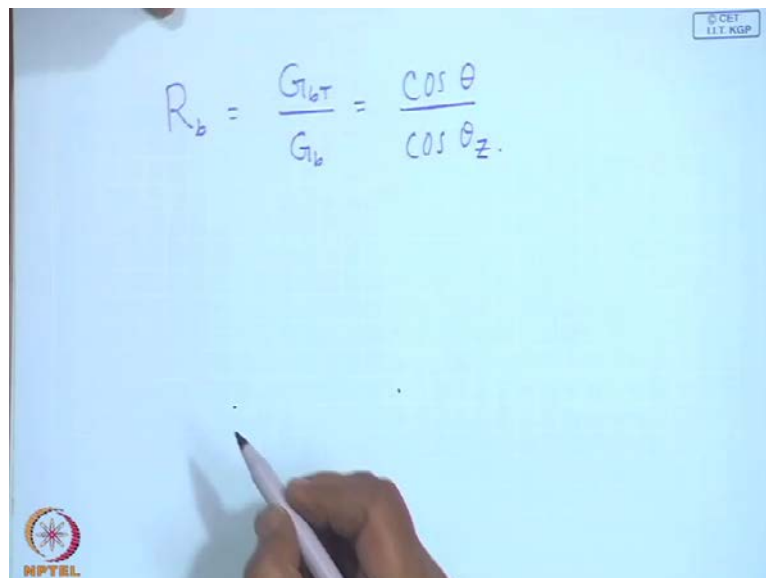
perpendicular to the sun's ray this is G_{bn} according to our relation G_{on} was a corresponding extraterrestrial value.

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Now, I will call it G_{bn} the still value of the direct radiation normal to the sun's ray, that, also will be G_{bn} the reason being this is the plane normal to the sun's ray. Whether there is a tilted surface or a horizontal surface is no consequence. So, the horizontal radiation G_b will be either G_{bn} into $\cos \theta_z$ here projection of this, or G_{bn} into $\cos \theta$ which will be G_{bT} .

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$$R_b = G_{bT} / G_b = \cos \theta / \cos \theta_z$$

Recall the expression for $\cos \theta$ and $\cos \theta_z$

R_b for a south facing surface is given by,

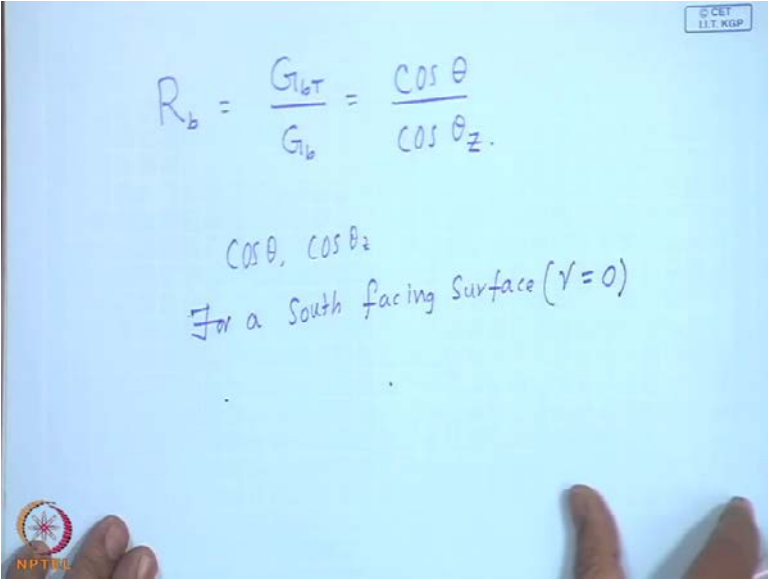
$$R_b = \frac{\cos(\phi - \beta) \cos \delta \cos \omega + \sin(\phi - \beta) \sin \delta}{\cos \phi \cos \delta \cos \omega + \sin \phi \sin \delta}$$

Strictly speaking is valid for an instant designated by the hour angle ω
However, the above can be used with no significant loss of accuracy for a time period of one hour (ω to $\omega + \omega/2$).

Thus, the overall tilt factor R , defined as, G_r/G can be expressed as,

$$R = \frac{G_b}{G} R_b + \frac{G_d}{G} \left(\frac{1 + \cos \beta}{2} \right) + \left(\frac{1 - \cos \beta}{2} \right) \rho_x$$


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$$R_b = \frac{G_{bT}}{G_b} = \frac{\cos \theta}{\cos \theta_z}$$

$\cos \theta, \cos \theta_z$
 For a South facing surface ($\gamma = 0$)

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So, now R_b is by definition. G_{bT} by G_b equal to $\cos \theta / \cos \theta_z$. Now, we have got this expression for R_b . And you can recall your $\cos \theta$ expressions and $\cos \theta_z$ expressions. And I can remember easily for a south facing surface that is your surface at the tilt angle is $\gamma = 0$.

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$$R_b = \frac{\cos(\phi - \beta) \cos \delta \cos \omega + \sin(\phi - \beta) \sin \delta}{\cos \phi \cos \delta \cos \omega + \sin \phi \sin \delta}$$

~~R = G~~

$$G_T R = G_T$$

So, now my R_b can be written as easy way of remembering is this is $\cos \theta_z$ which will not depend upon γ . This is $\cos \theta_z$ for γ equal to 0. So, if you put a β equal to 0 you will get this result. So, this is R_b for a south facing surface. Now, what I shall do is, if I define a overall R factor. Which is G_b / G_T let me put it this way G_T total tilted radiation will be your G into R then that will be equal to G_T .

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$$R = \frac{G_b}{G} R_b + \frac{G_d}{G} \left[\frac{1 + \cos \beta}{2} \right] + \rho_g \left[\frac{1 - \cos \beta}{2} \right]$$

$$= \frac{(G - G_d)}{G} R_b + \frac{G_d}{G} \left[\frac{1 + \cos \beta}{2} \right] + \rho_g \left[\frac{1 - \cos \beta}{2} \right]$$

$$\left(1 - \frac{G_d}{G} \right) R_b + \frac{G_d}{G} \left[\frac{1 + \cos \beta}{2} \right] + \rho_g \left[\frac{1 - \cos \beta}{2} \right]$$

So, now I can define a, overall r factor as the same thing I have written. If we multiply throughout by G into R will be your G_T . And hence, this is written in a particular


fraction. Now, you will find that everywhere I have got a G_b by G , G_d by G . Of course, this G gets canceled this can be again written as G minus G_d by G times R_b plus G_d by G times $1 + \cos \beta$ by 2 plus ρ_g G $1 - \cos \beta$ by 2 . This is $1 - \rho_g$ by G .

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The hourly tilted radiation I_T calculating R_b at $\omega = (\omega_1 + \omega_2)/2$ can be expressed as,

$$I_T = RI = I_b R_b + I_d \left(\frac{1 + \cos \beta}{2} \right) + \rho_g I \left(\frac{1 - \cos \beta}{2} \right)$$



Expressing $I_b = I - I_d$, R for a small period of time can be evaluated from,

$$R = \left(1 - \frac{I_d}{I} \right) R_b + \frac{I_d}{I} \left(\frac{1 + \cos \beta}{2} \right) + \rho_g \left(\frac{1 - \cos \beta}{2} \right)$$


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$$R = \frac{G_b}{G} R_b + \frac{G_d}{G} \left[\frac{1 + \cos \beta}{2} \right] + \rho_g \left[\frac{1 - \cos \beta}{2} \right]$$

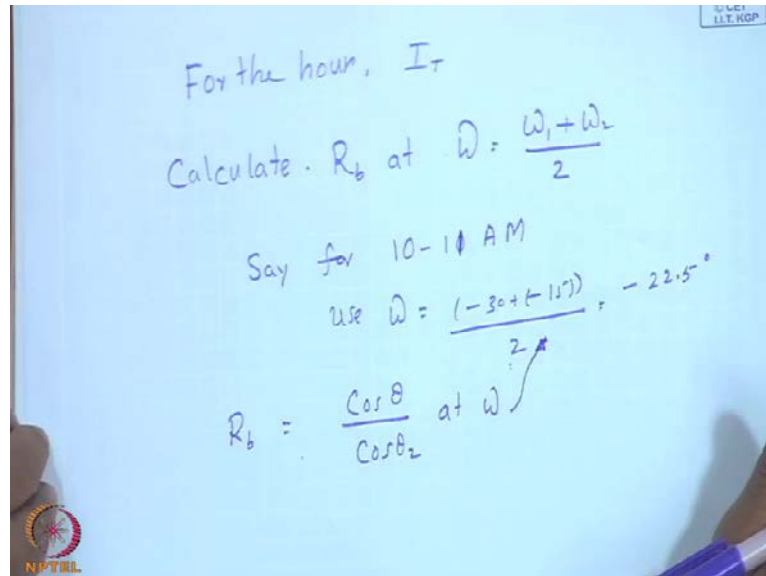
$$= \frac{(G - G_d)}{G} R_b + \frac{G_d}{G} \left[\frac{1 + \cos \beta}{2} \right] + \rho_g \left[\frac{1 - \cos \beta}{2} \right]$$

$$\left(1 - \frac{G_d}{G} \right) R_b + \frac{G_d}{G} \left[\frac{1 + \cos \beta}{2} \right] + \rho_g \left[\frac{1 - \cos \beta}{2} \right]$$



Now, what I am finding out is that this is a ratio of only G_d by g , now you realize why we spend some time in generating the correlations for i_d by I . Though I_d . And I are

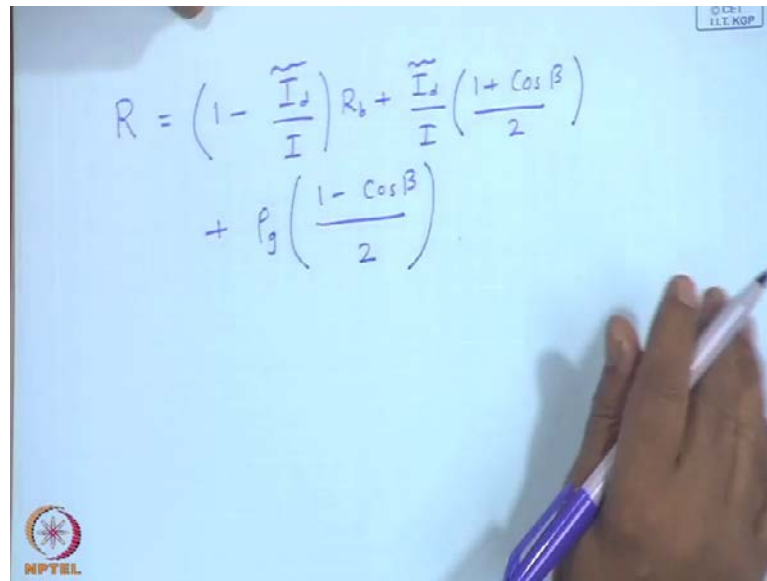
separately required. But, if given I everything we try to write in terms of a, diffused fraction ratio which we know can be calculated knowing I.

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Now, I shall make an approximation for the hour the corresponding value will be I_T right instead of G_t it will be I_T calculate R_b at mid 1 midpoint of the hour. That is say for 10 to 11 use omega is equal to, How much it will be this is minus 15 minus 30 correct and calculate your R_b . Which in general simply $\cos \theta$ by $\cos \theta_z$ at this omega given by $\omega_1 + \omega_2$ by 2..

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A hand is writing the following equation on a whiteboard:

$$R = \left(1 - \frac{\tilde{I}_d}{I}\right) R_b + \frac{\tilde{I}_d}{I} \left(\frac{1 + \cos \beta}{2}\right) + \rho_g \left(\frac{1 - \cos \beta}{2}\right)$$

The whiteboard also features the NPTEL logo in the bottom left corner and a small copyright notice in the top right corner.

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This form of the equation is valid for a non-south facing surface as well, when R_b is evaluated using

$$\cos \theta = A + B \cos \omega + C \sin \omega$$

where

$$A = \sin \delta (\sin \phi \cos \beta - \cos \phi \sin \beta \cos \gamma)$$
$$B = \cos \delta (\cos \phi \cos \beta + \sin \phi \sin \beta \cos \gamma)$$
$$C = \cos \delta \sin \beta \sin \gamma$$

The slide also features the NPTEL logo in the bottom left corner.

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$$R = \left(1 - \frac{\tilde{I}_d}{I}\right) R_b + \frac{\tilde{I}_d}{I} \left(\frac{1 + \cos \beta}{2}\right) + P_g \left(\frac{1 - \cos \beta}{2}\right)$$

For $\gamma \neq 0$

$$\cos \theta = A + B \cos \omega + C \sin \omega$$

A, B, C are defined.

If, I straight away extend the result that we have written for the instantaneous values. my factor R will be 1 minus I d by I, times Rb plus Id by I, 1 plus cos beta by 2 plus rho G1 minus cos beta by 2. Now, we got everything in terms of Id by I. So, you can easily write for we have this we have this general expression. A plus b cos omega plus C sine, omega where A,B,C are defined. So, fortunately these factors.


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F_s and F_g do not depend on γ or ϕ or ω or δ

$$R_b = \frac{A + B \cos \omega + C \sin \omega}{\cos \phi \cos \delta \cos \omega + \sin \phi \sin \delta}$$

F_s and F_g do not depend on gamma or phi or omega or even delta. So, I can in general have this is the general expression for cos theta. And this expression for cos theta z.

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SUMMARY

It is realized that the solar radiation on a tilted surface comprises of , beam, sky diffuse and ground reflected components.

It is assumed that the diffuse sky radiation is isotropic.

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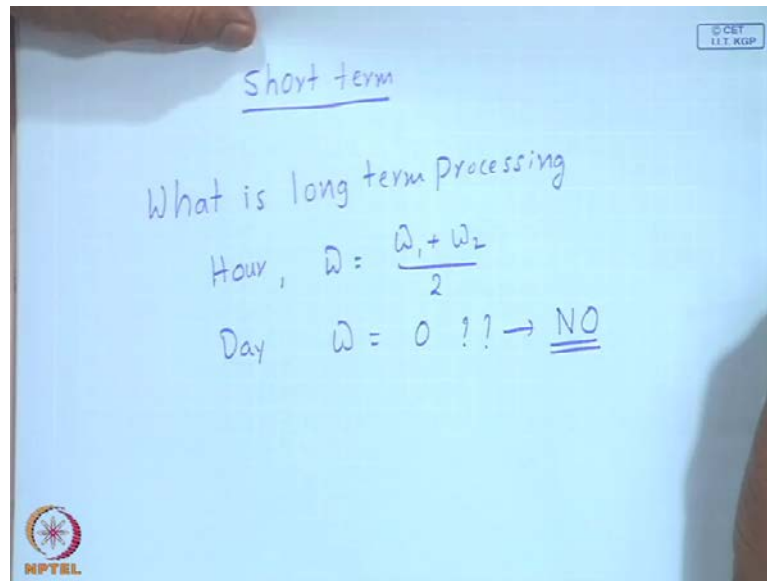
The tilted surface receives a portion of the diffuse and ground reflected radiation. The factors F_s and F_g being functions of the slope of the surface only.

Though the factor, R_b , *strictly speaking*, is valid for an instant, can be applied for a short period of time interval, say, up to 1 hr,

by evaluating R_b at the mid point of the hour.

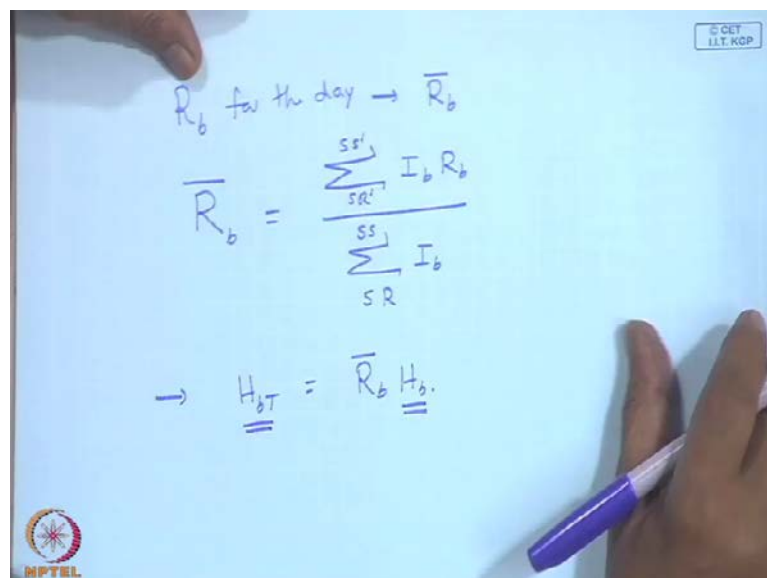
So, now we have a method of calculating r_b in general. If we summarize solar radiation on a tilted surface comprised of beam sky diffuse and ground reflected components. And it is assumed that this sky diffuse radiation is isotropic, and the tilted surface this is a portion of the sky diffuse and the ground reflected radiations. And though the R_b factor is strictly to be applicable to be calculate for an instance without significant or almost 0 loss of accuracy. You can calculate at the midpoint of the hour. Now this is short term processing.

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We will have little bit of introduction for what is long term. For an hour we are able to calculate ω_1 plus ω_2 by 2, What about a day can I calculate the midpoint of the day, which is a solar room which is ω is equal to 0? Answer is obviously a big no because, it will have the if it is a south facing surface. You will have the perhaps best orientation at the noon time. And hence, R_b calculated ω is equal to 0 cannot be represented of the entire day.

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But, how do I define if I want to have a daily R_b for the day, let us call it \bar{R}_b like we have been doing it. So, this \bar{R}_b should be equal to summation of $I_b R_b$ summation of I_b from again categorically from sunrise to sunset sunrise to sunset.

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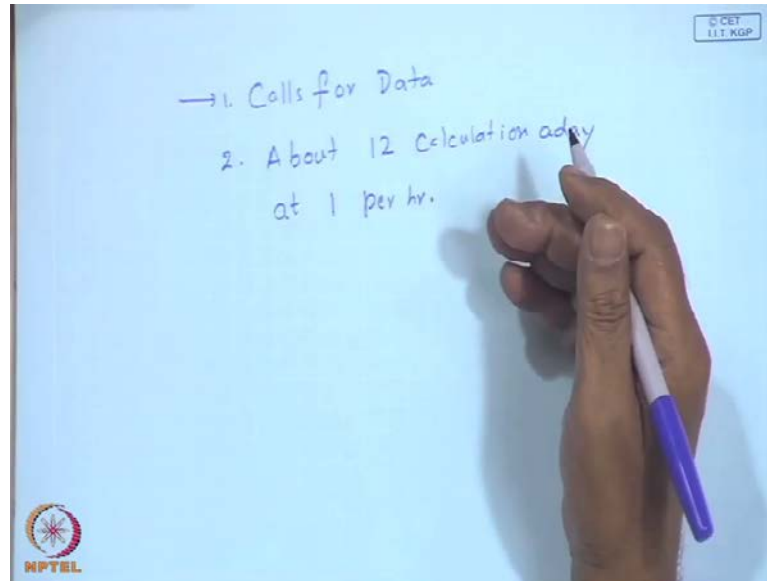
\bar{R}_b → is a Weighted Avg.

Re-write in terms of G

$$\bar{R}_b = \frac{\int_{SS'}^{SR'} G_b R_b dt}{\int_{SR}^{SS} G_b dt}$$

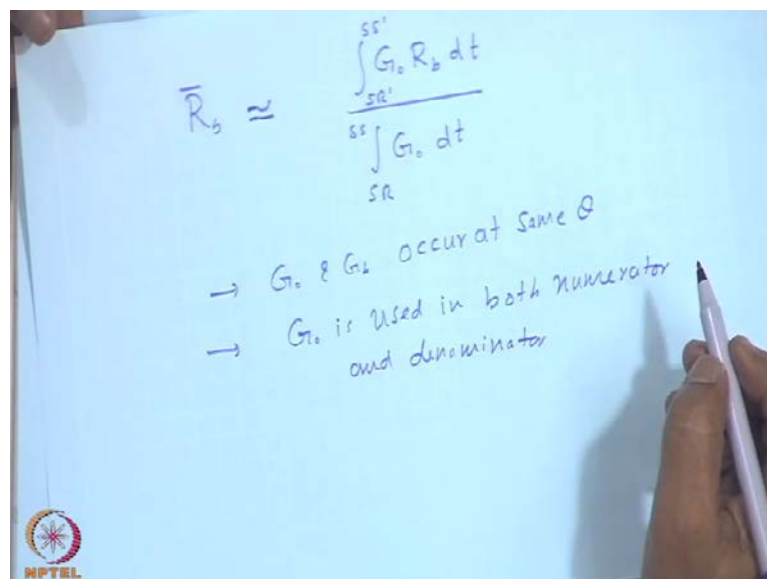
Little later we will find out these are not the physical sunrise and sunset but, this should be sum SR dashed. And SS dashed which apparently the surface sees during the day. But, never the less if we properly calculate it even if I categorically say from sunrise to sunset and ignore a negative R_b etcetera. You will still get a correct value. So, if I am asked to find out \bar{R}_b or in general this leads to h_t should be equal to $\bar{R}_b h_b$.

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If, I have the daily beam radiation on a horizontal surface, I will be able to calculate the daily direct radiation on the tilted surface. By this factor \bar{R}_b into h_b . so, this is \bar{R}_b is a weighted average assuming I can transfer just to be mathematically consistent. Either I sum up or by over R_f calculate for categorical reasons as $G_b R_b dt$ in terms of an integral.

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But, this calls for data 1, 2 about 12 calculations a day at 1 per hour. all the days may not exactly have 12 hours. But, about that on an yearly basis this average is. so this is a

conversion process. And if you want to calculate for the month you need to calculate for 12 into 30 days for the year it will be even more multiplied by 365. So, how to solve this problem what people have suggested is the data calls for the numbers and a numerical integration. But if you approximately calculate under extraterrestrial conditions, what are my justifications justifications are G_0 . And G_b occur at same theta, the angle of incidence of the extraterrestrial radiation on a horizontal plane.

And the instantaneous G_b at a instant will have same theta what differs is G_0 numerically differs from G_b . So this is an approximate symbol based upon this then second sort of justification is a G_0 is used in both numerator and denominator. So, if there is an error in the numerator there is some error in the denominator. Also though not exactly the same consequently some error is likely to get canceled. So, this appears to be a good approximation.

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In terms of ω , the hour angle

$$\bar{R}_b \approx \frac{\int_{\omega_{SR}}^{\omega_{SS}} G_{0m} \cos \theta_z \frac{\cos \theta}{\cos \theta_z} d\omega}{\int_{-\omega_S}^{\omega_S} G_{0m} \cos \theta_z d\omega}$$

The numerator is simplified to $G_0 \int_{\omega_{SR}}^{\omega_{SS}} \cos \theta d\omega$ and the denominator is $G_0 \int_{-\omega_S}^{\omega_S} \cos \theta_z d\omega$.

Converting into in terms of omega the hour angle bar approximately, I will now call it omega Sr to omega SS and this G_0 will be $\cos \theta$ times $\cos \theta_z$. So, this is a normal radiation multiplied by $\cos \theta_z$ which will be the nothing but, G_0 which is a same G_0 over here. and my R_b is categorically $\cos \theta$ upon $\cos \theta_z$. I am now distinguishing the tilted surface may not see the radiation all the time. As the physical sun sees the horizontal surface or above the horizon.

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The image shows a handwritten derivation on a whiteboard. At the top right, there is a small box containing the text "© CET IIT KGP". The derivation starts with the equation:

$$\bar{R}_b \approx \frac{\int_{\omega_{SR}}^{\omega_{SS}} G_{\text{sun}} \cos \theta_z \frac{\cos \theta}{\cos \theta_z} d\omega}{\int_{-\omega_s}^{\omega_s} G_{\text{sun}} \cos \theta_z d\omega}$$

Below this, it is simplified to:

$$\bar{R}_b \approx \frac{\int_{\omega_{SR}}^{\omega_{SS}} \cos \theta d\omega}{\int_{-\omega_s}^{\omega_s} \cos \theta_z d\omega}$$

The NPTEL logo is visible in the bottom left corner of the whiteboard image.

Let me rewrite it at the moment do not bother about this omega s r and omega s s except physically except that any tilted surface are arbitrarily oriented surface. We, shall not see the sun all the time that the sun is above the horizon. So, this simplifies to you can write GSC into $1 + 0.033 \cos 360 n \text{ by } 365$. And this $\cos \theta_z$ gets cancelled with this $\cos \theta_z$. And this also cancels with this so, my \bar{R}_b is approximately equal to integral omega SR to omega SS of $\cos \theta d\omega$ by minus omega S to plus omega S $\cos \theta_z d\omega$.

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The image shows handwritten notes on a whiteboard. At the top right, there is a small box containing the text "© CET IIT KGP". The notes state:

\bar{R}_b is NOT very complicated

\bar{R}_b fairly Accurate against \bar{R}_h With Data

$$\frac{\sum I_b R_b}{\sum I_b}$$

More Accurate for $\gamma = 0$.

The NPTEL logo is visible in the bottom left corner of the whiteboard image.

So, \bar{R}_b is not very complicated. So, surprisingly it looks like a that you are integrating a numerator of R_b . And the denominator of R_b separately that is integral $\cos \theta$ by integral $\cos \theta z$, to get \bar{R}_b though that is not a mathematically correct argument. And you it so, turned out that $\cos \theta z$ term gets canceled. And hence you will have this result. And this \bar{R}_b fairly accurate against R_b with data, in other words you compute this with $\sum I_b R_b$ by $\sum I_b$. This will be within 3 to 5 percent and this is more accurate γ is equal to 0.

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$$G_b \approx G_0$$

→ Calculate Under ET Condition
Why? accurate.

$$\bar{R}_b = \frac{\int I_b R_b dt}{\int I_b dt}$$

$$= \frac{\int (I - I_d) R_b dt}{\int I_b dt}$$

So, can we find any reason why this works well. Now, many people have made calculations we also made calculations, we found this \bar{R}_b obtained. So, called extraterrestrial conditions basically G_b being equated to G_0 . That means calculate under Et now we shall use this short form for extraterrestrial condition. Now you may think of, Why, we can put forth 1 argument again? If we do not bother to much about the rate. And the intensity or the hourly value, basically it is $I_b R_b dt$ by $I_b dt$ limits are un important for the argument that is going to follow. So, this is I minus $I_d R_b dt$.

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$$\bar{R}_b = \frac{\int I \left(1 - \frac{I_d}{I}\right) R_b dt}{\int I \left(1 - \frac{I_d}{I}\right) dt}$$
$$= \frac{\int k_T I_0 \left(1 - \frac{I_d}{I}\right) R_b dt}{\int k_T I_0 \left(1 - \frac{I_d}{I}\right) dt}$$

I am trying to assume that I have got analytically representable function, or discredited function or some correlation for I. And I write it in this particular fashion this is exact under terrestrial conditions. Now we this once again so, what all I have done is the terrestrial radiation is nothing. But, clearness index multiplied by the extraterrestrial radiation this everything else is now uniform and we realize a sum function of clearness index.

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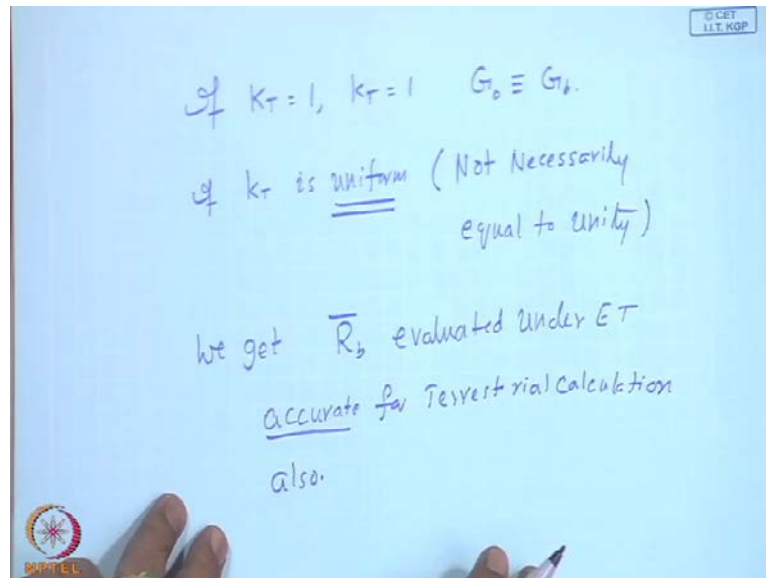
Now if k_T is uniform and
we realize $\left(1 - \frac{I_d}{I}\right) = f(k_T)$

$$\rightarrow \bar{R}_b = \frac{\int I_0 R_b dt}{\int I_0 dt} \rightarrow \text{ET Calculation}$$

You can rewrite in terms of G_0

Because, all I_d by I correlations and h_d by h correlations are just functions of K_T mainly. If that is so, my \bar{R}_b will be simply you have $I_0 R_b dt$ by $I_0 dt$ that means ET calculation. You can again I can make it the intensity and bring back my old theory.

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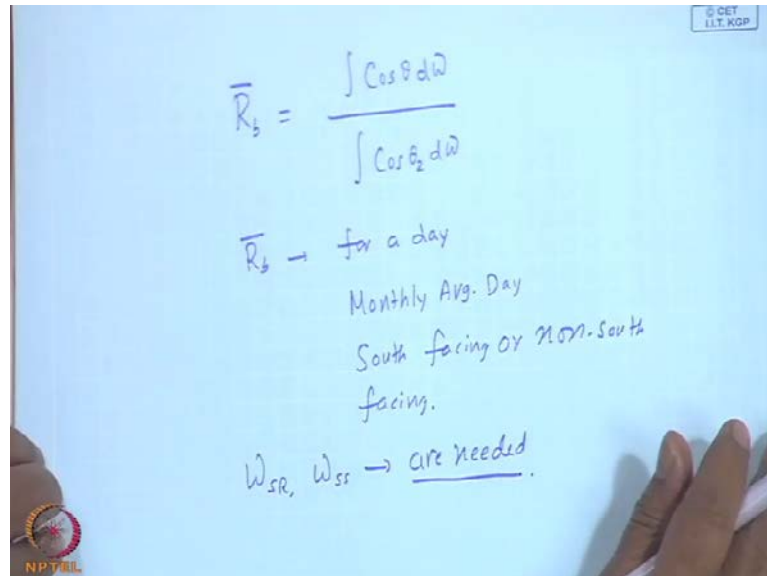


Now, what did we achieve by showing this what we achieved was what we realized is, if the clearness index has been 1. We, know that extraterrestrial and terrestrial radiation would be the same. There will be no diffused radiation by this we proved if K_T small K_T is uniform. That means not necessarily equal to unity, we get \bar{R}_b evaluated ET accurate enough for terrestrial calculation also. What really in simple English means is you need not have atmosphere transmittance unity. But, you need to have uniform atmosphere transmittance.

That means if you have a uniformly cloudy or uniformly bright day your extraterrestrial calculation will be very close to the terrestrial calculation. And there are little favorable things if you think little more 80 percent of the radiation will be about, 5 to 6 hours mid part of the day 10 to 4. For example, that is when your R_b factors will be high that is also when your clearness index is likely to be more or less uniform. So, the weightage because of non-uniformity occurring at a lower radiation values will be small there. By making \bar{R}_b calculated under extraterrestrial conditions. With a analytical expression will be fairly close to your values that you will have obtained using the data. Of course,

there are deviations when the surface is not to south under certain extreme conditions. But, most of the time the extraterrestrial approximate calculation is acceptable.

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So, \bar{R}_b is simple before we proceed to examine this in fact even \bar{R}_b this can be for a day then a monthly average day. And it could be south facing or non south facing. So, we shall try to evaluate this occurred ω_{SR} and ω_{SS} are needed. So we have a method of calculating the tilted radiation for a hour, we have it for a day which works well. And we will examine how it performs for the monthly average day. And how we estimate this so, called apparent sunrise and sunset hour angles for the tilted surface.